

Multi-beam SIS Receiver Development

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Still

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List of Acronyms

- IRAM: Institut de RadioAstronomie Millimétrique
- RF: Radio Frequency
- IF: Intermediate Frequency
- SIS: Supraconducting-Isolator-Supraconducting
- SSB: Single Side Band
- 2SB: Two Side Bands
- LNA: Low Noise Amplifiers
- OMT: Ortho Mode Transducer
- GM-JT: Gifford-Macmahon-Joule-Thomson.
- GM: Gifford-Macmahon
- FOV: Field Of View
- HPBW: Half Power Beam Width
- PTFE: Poly Tetra Fluoro Ethylene
- HDPE: High Density Poly Ethylene



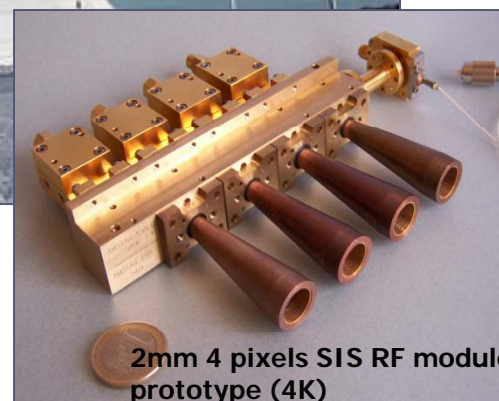
IRAM Pico Veleta Telescope



Altitude: 2850 m
Localisation: Sierra Nevada
(Andalusia, Spain)

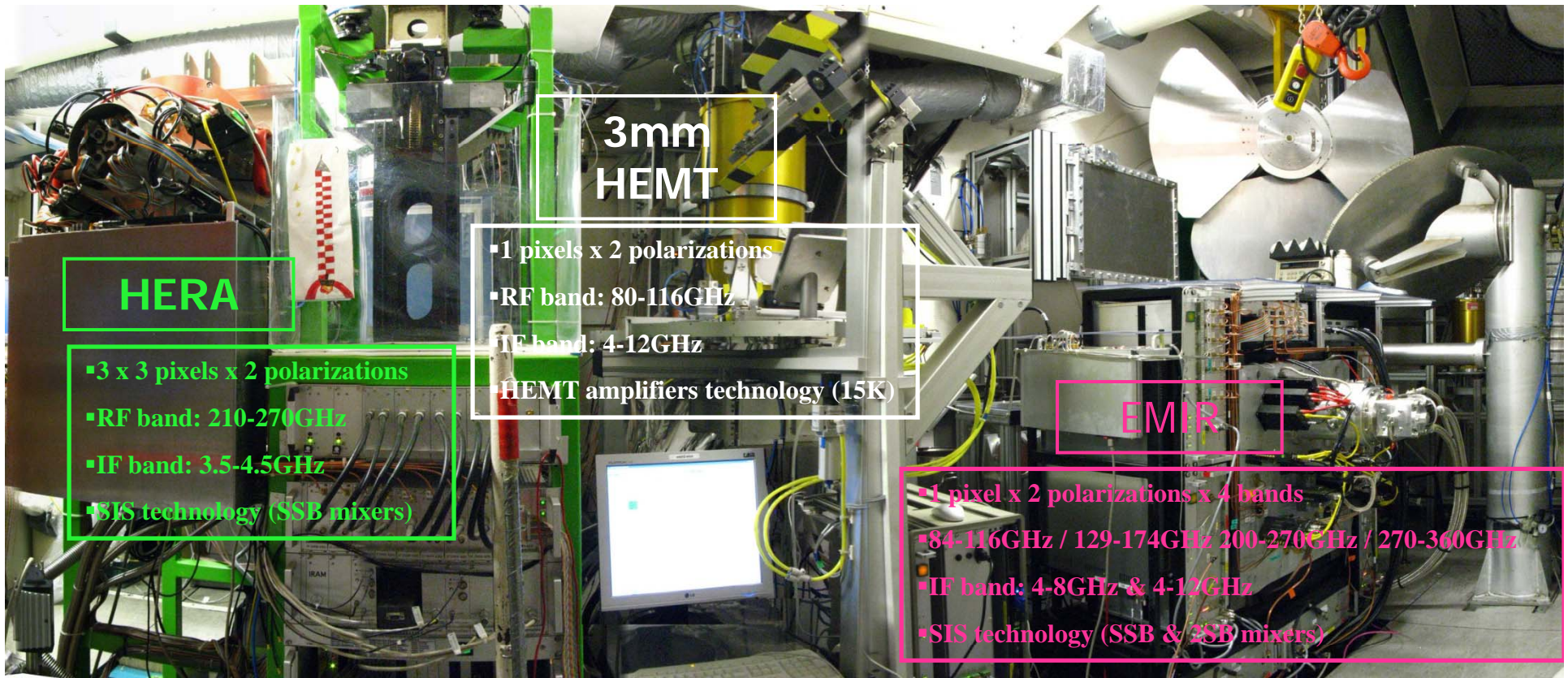
Interest in developing Multi-beam SIS Receivers:

- Increase mapping speed of extended sources
- Improve data quality





Current Heterodyne Instrumentation @ Pico Veleta



HERA

- 3 x 3 pixels x 2 polarizations
- RF band: 210-270GHz
- IF band: 3.5-4.5GHz
- SIS technology (SSB mixers)

**3mm
HEMT**

- 1 pixels x 2 polarizations
- RF band: 80-116GHz
- IF band: 4-12GHz
- HEMT amplifiers technology (15K)

EMIR

- 1 pixel x 2 polarizations x 4 bands
- 84-116GHz / 129-174GHz 200-270GHz / 270-360GHz
- IF band: 4-8GHz & 4-12GHz
- SIS technology (SSB & SSB mixers)

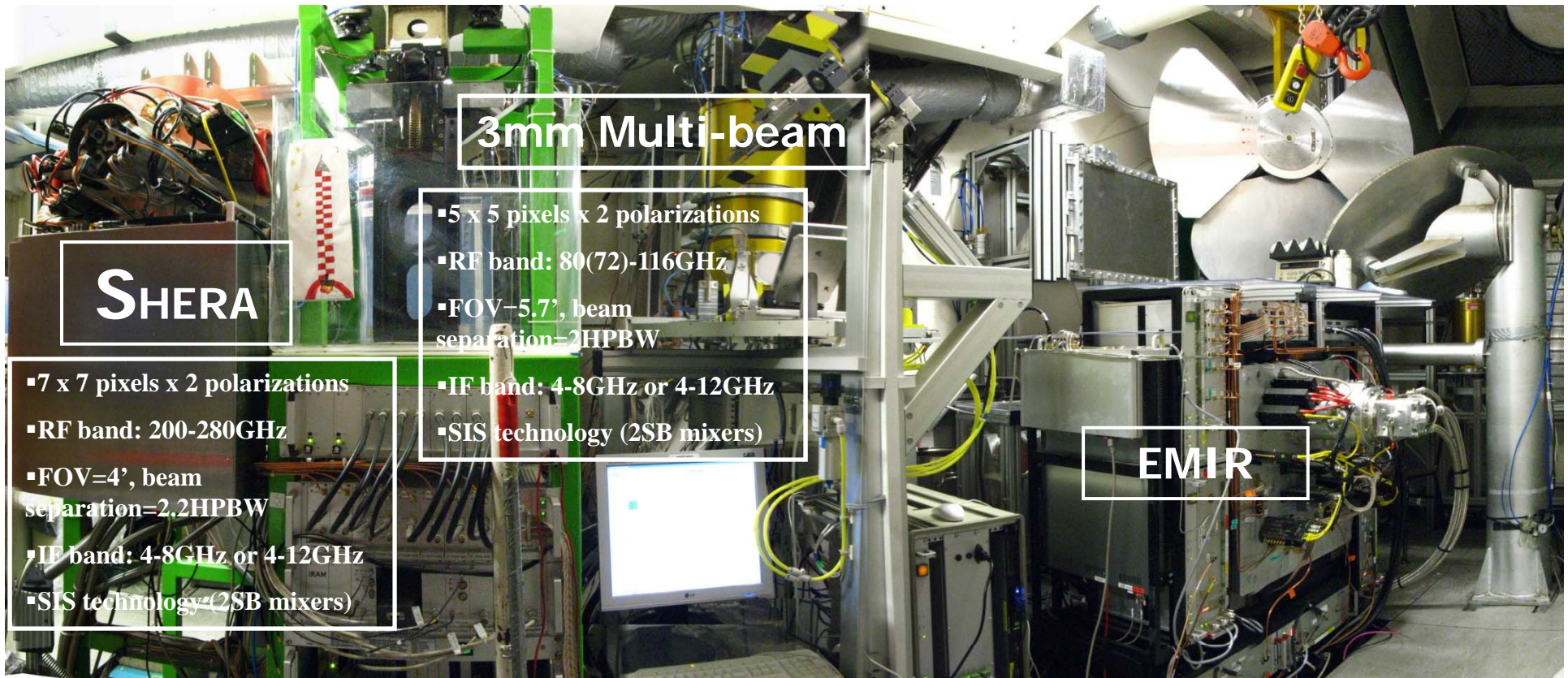
GM-JT Daikin cryocooler CG308SC
3 stages: 77K /15K /4K(2.5W)
Coldest T° ~ 4.2 K (unloaded stage)

HDV10 cryostat
2 stages: 77K/ 15K

GM Sumitomo cryocooler RDK-3ST
3 stages: 77K /15K /4K(1W)
Coldest T° ~ 3 K (unloaded stage)

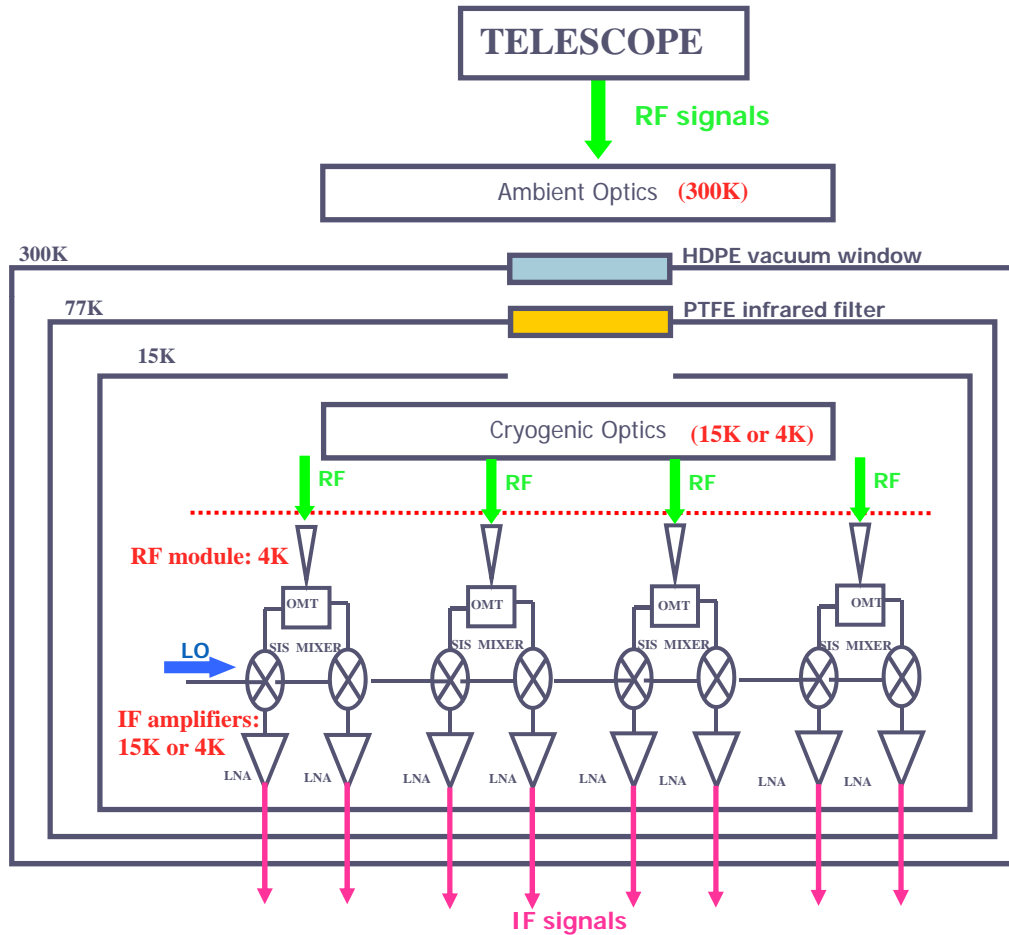


Future Heterodyne Instrumentation @ Pico Veleta

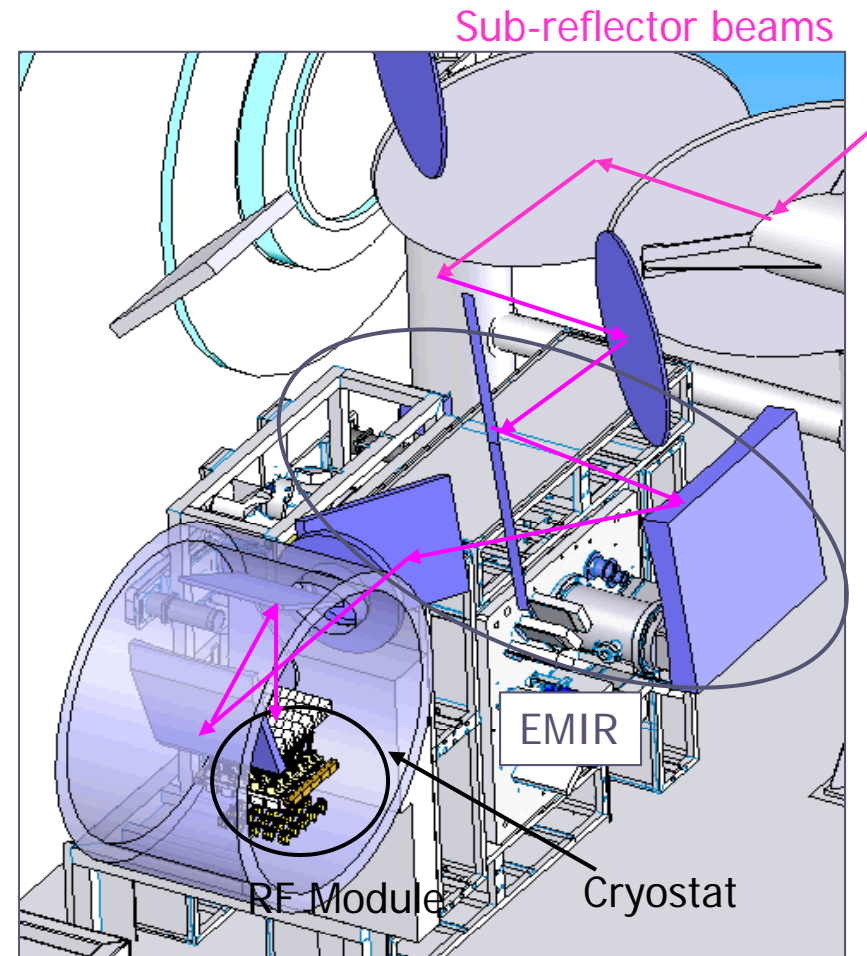


Multi-beam SIS Receiver Overview

Synoptic of a 4-pixel array



3D view of the future 3mm multi-beam



Requirements and Critical Design Items

Main requirements:

- Compact size
- Easy to be repaired or upgraded
- State of the art performances (noise, stability, optics)

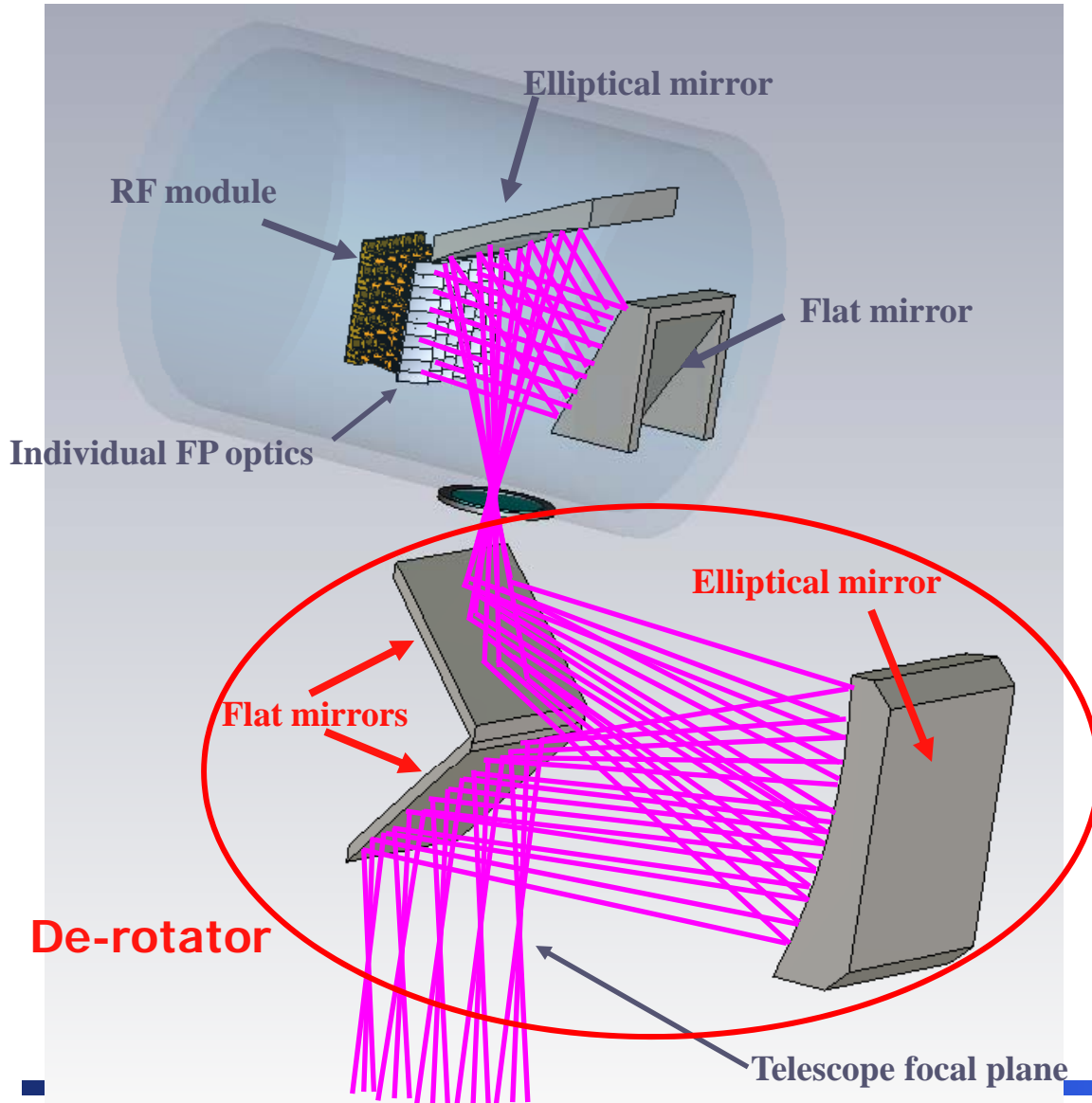


Some critical points in the receiver design:

- Optics design → receiver size, cryostat size, receivers performances...
- RF module design → cryostat size, receiver performances, repairing procedure
- Cryogenic aspects → cryostat size, receiver performances, repairing procedure



Optics Design Overview of the 3mm Multibeam



- Frequency independent sub reflector illumination (taper = -12dB)

- Transform 2HPBW (=68.5mm in FP) spacing on the sky into 42mm between feeds

- Limited thermal radiation due to 300K window and IR filter

- 4W on 77K stage

- 80mW on 15K stage

- Optimal beam coupling between telescope and horn apertures

- K-mirror for image rotation

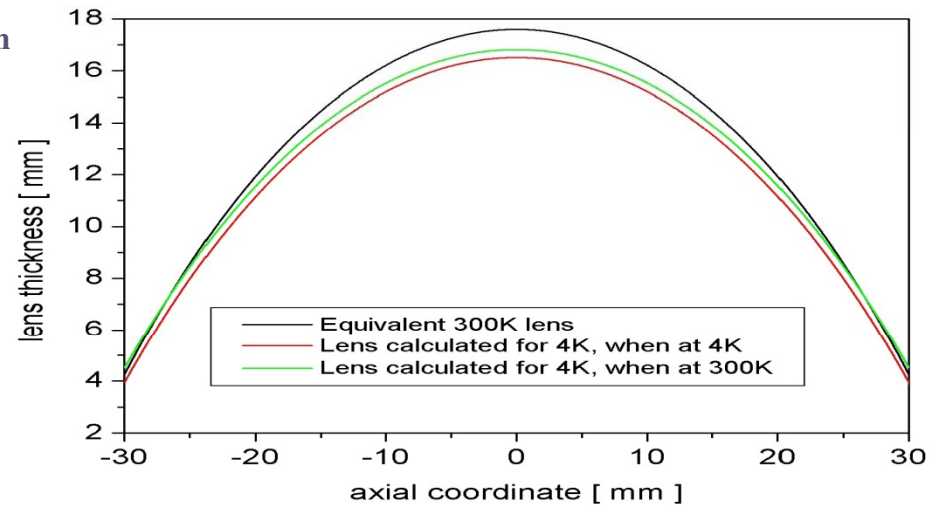
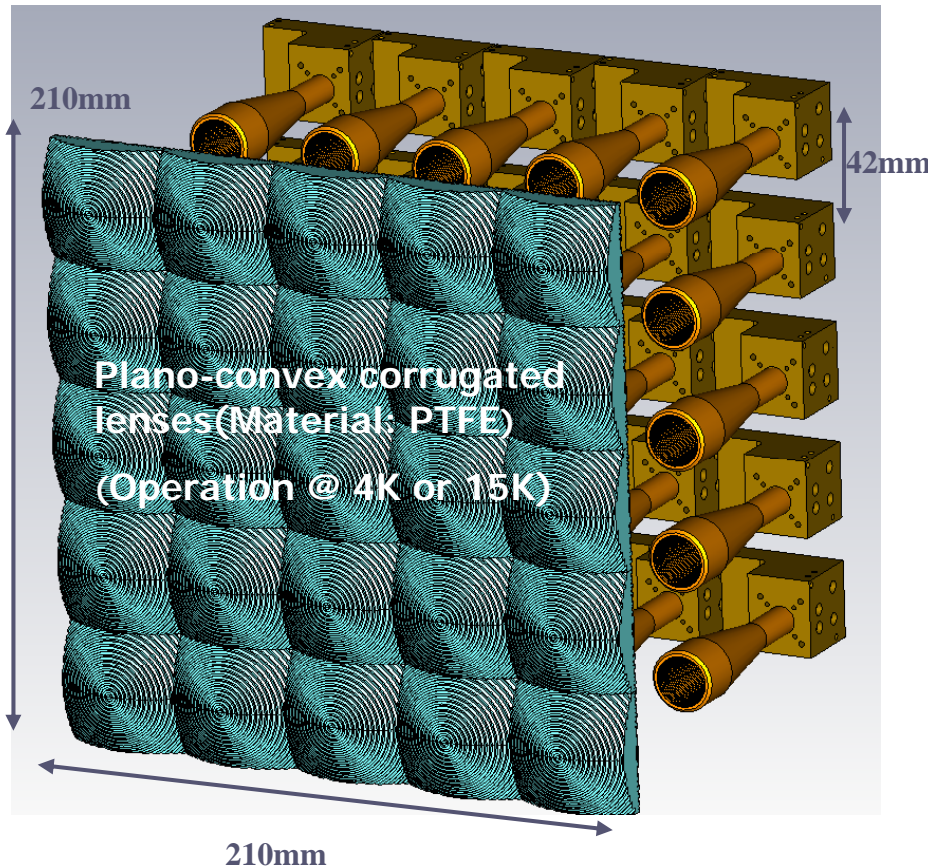


3mm MB: Individual Cryogenic Optics

Design Option A: Refractive Optics

Lens design @ 4K:

- Permittivity: $\epsilon_r = 2.07$ (300K) \rightarrow 2.16 (4K)
- Linear thermal contraction: 1.6% to 2.1%

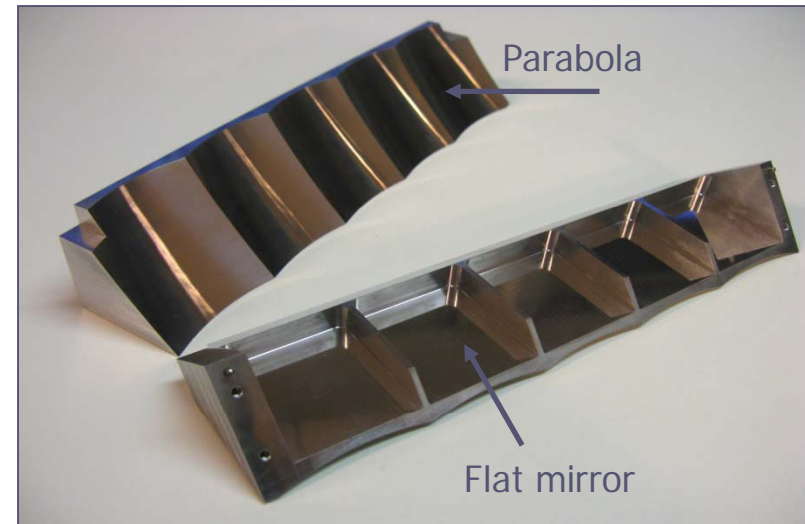
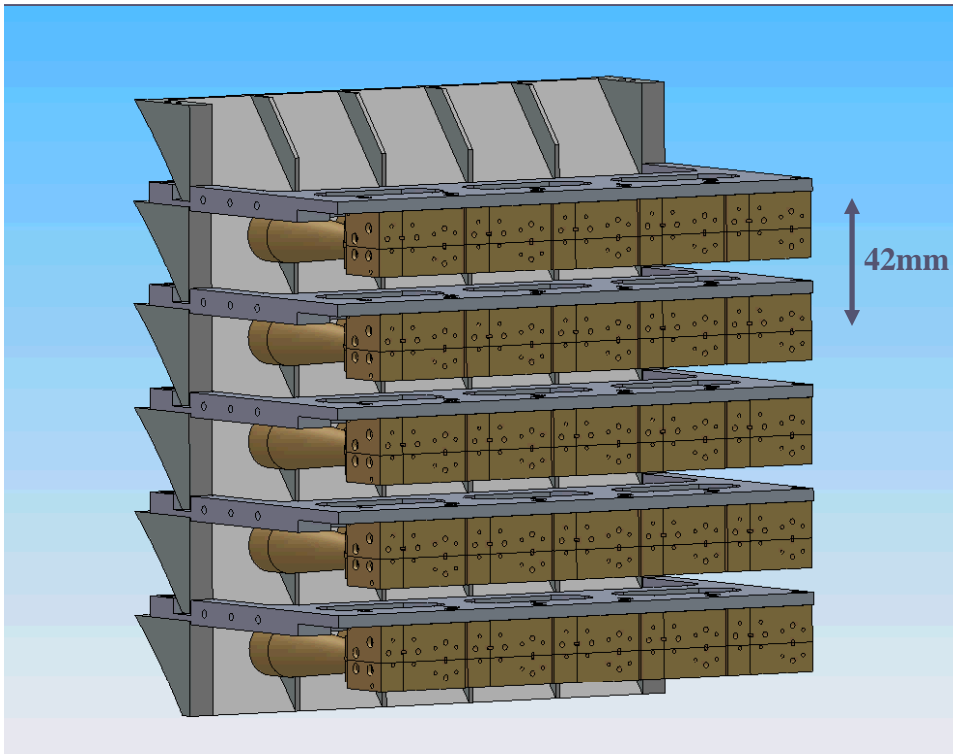


Effects of lenses absorption losses on receiver noise temperature:

\rightarrow ~ + 5K (@ 300K window output) if $T_{rec} = 50K$ (@ horn output)



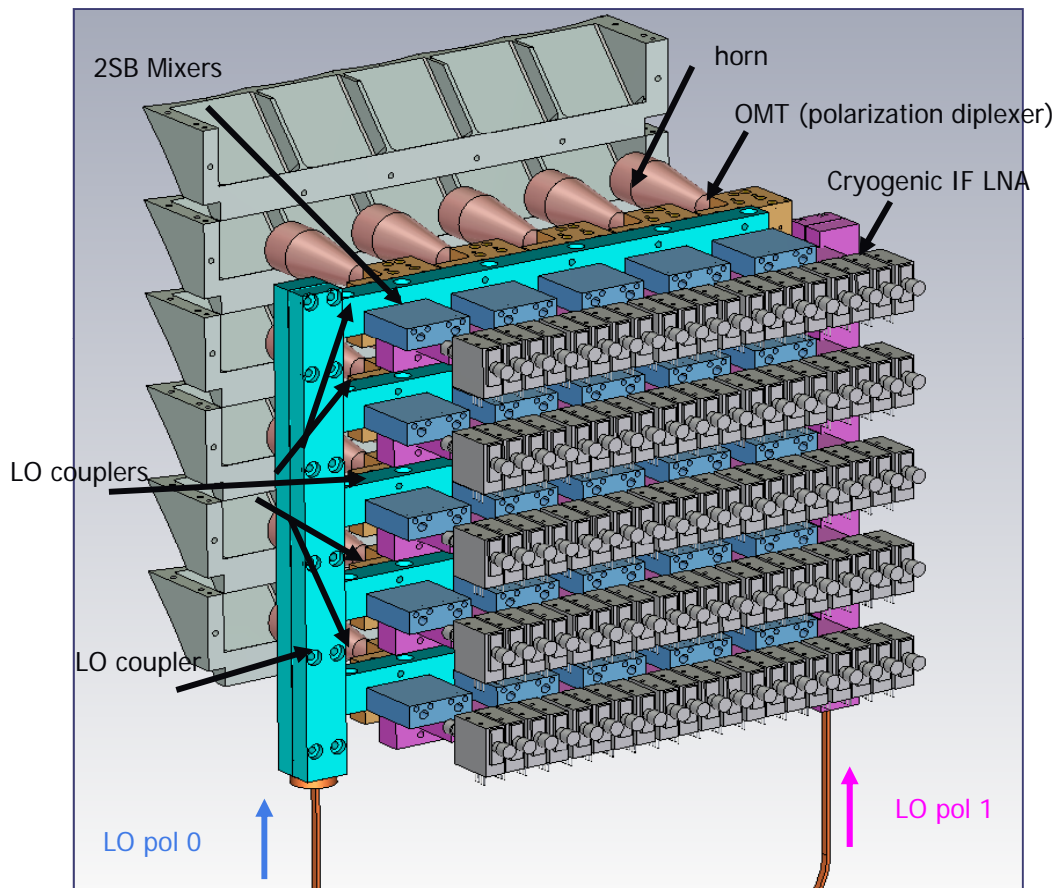
3mm MB: Individual Cryogenic Optics Design Option B: Fully Reflective Optics



- No absorption loss & reflection loss
- Negligible thermal contraction



3mm MB: Cryogenic RF Module



Thermal Issues:

- Manufacturing materials of cryogenic waveguide components :

Polarization diplexer (gold plated brass)



- Physical temperature of active components (mixers, LNA...)
- Dismounting/ repair procedure: cryogenic IF cables & connectors

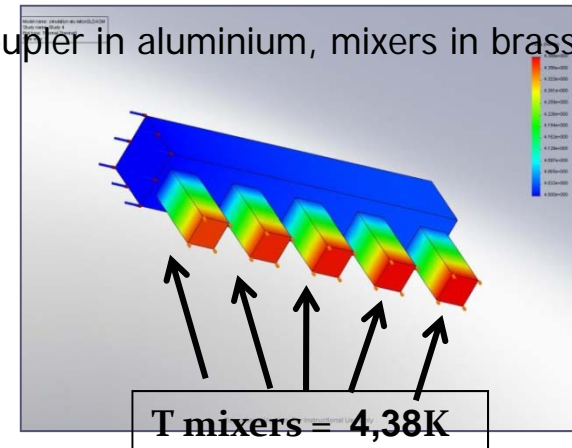
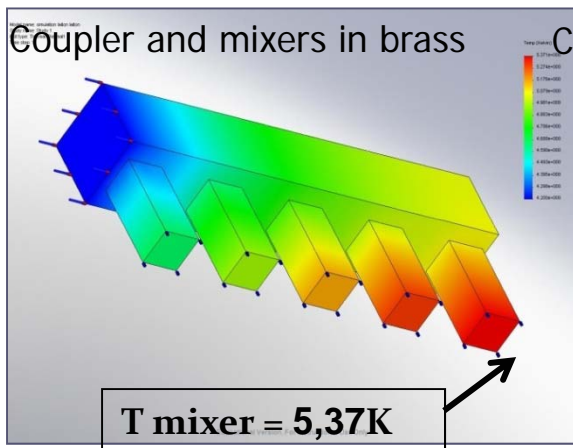
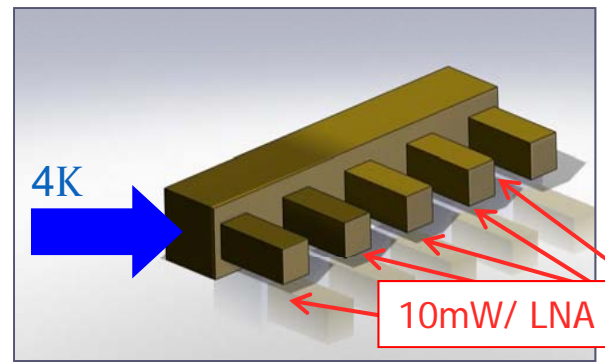
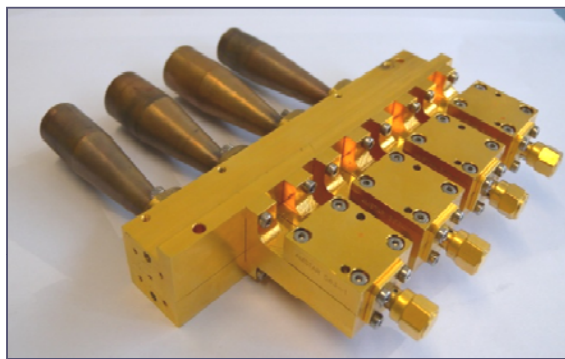


Cryogenic Waveguide Components: Materials Analysis

Material requirements:

- Low mass volume
- High electrical conductivity
- High thermal conductivity
- Accurate machining

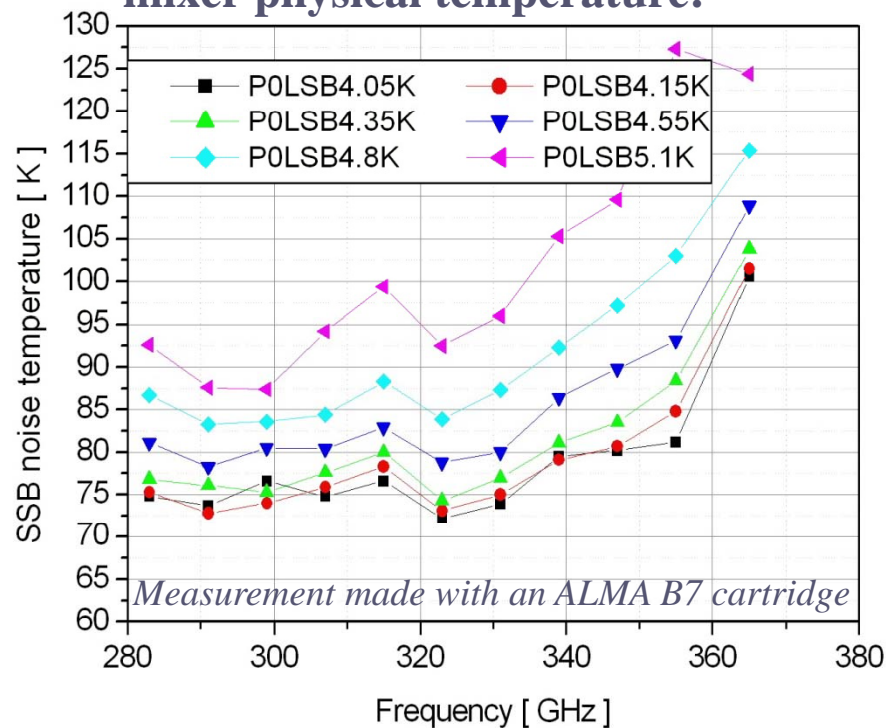
Thermal simulations



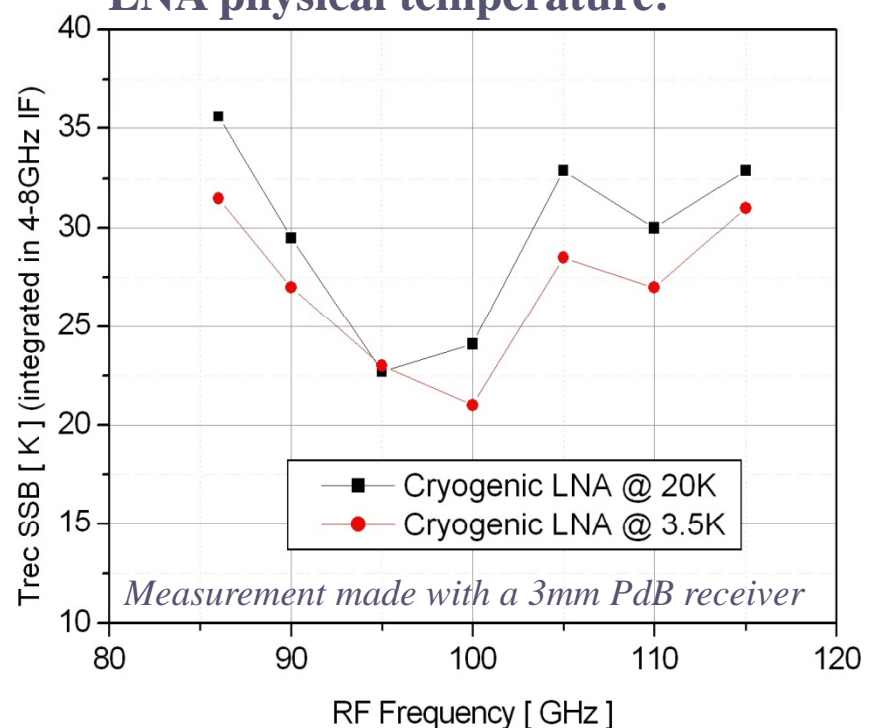


Cryogenic Design: Optimal Operating Temperatures of SIS Mixers and LNA

Receiver noise performances vs. SIS mixer physical temperature:



Receiver noise performances vs. LNA physical temperature:

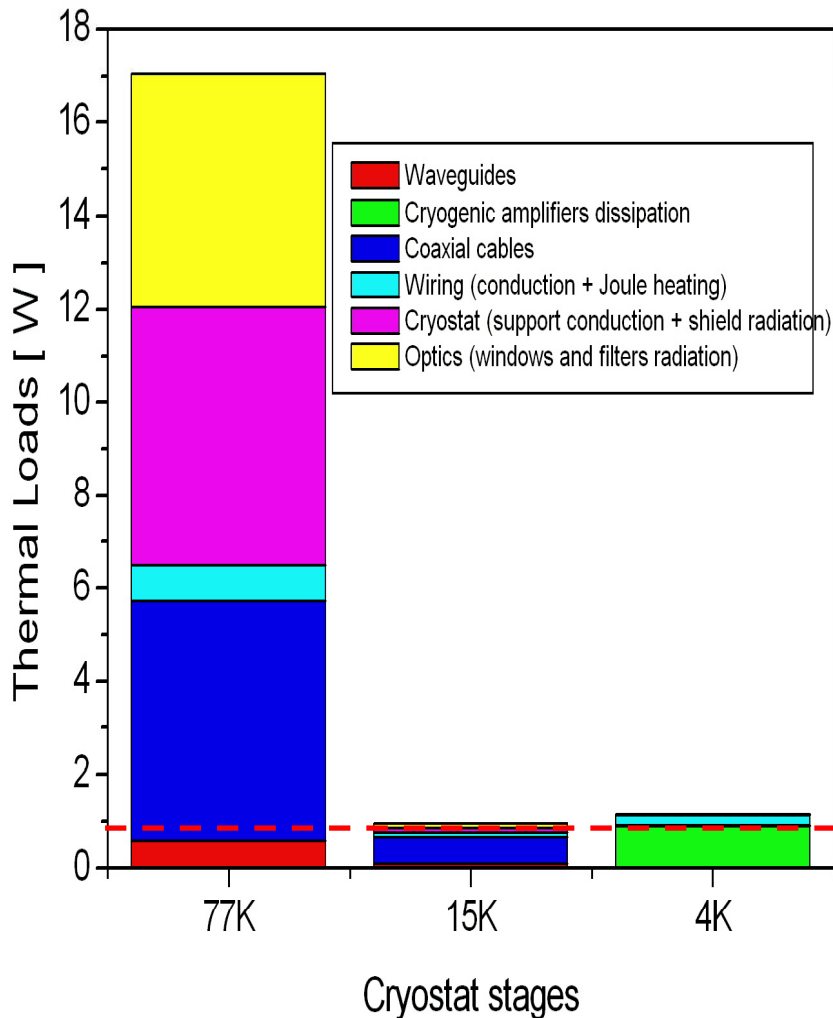


→ In the HERA cryostat (GM-JT DAIKIN cryocooler), mixers physical temperatures ~ 4.7K
 → In the EMIR cryostat (GM SUMITOMO cryocooler), mixers physical temperatures ~ 4K

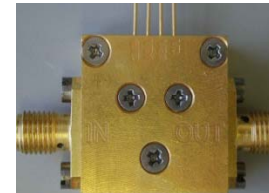


Cryogenic Design: Thermal Budget of the 3mm Multi-beam

(50 pixels, 2SB = 100 IF outputs are considered)



- Main contribution on 4K stage: LNA (9mW/ ampli)



CALTECH 4-12GHz amplifier

- IF transport: SS/CuBe semi-rigid cables (5W(77K)/560mW(15K)/20mW(4K))
(Other solutions are also considered for IF transport for mechanical reasons):



Flexible cryogenic cable from HIGHTEC

- Wires: MG, $\phi=0.2\text{mm}$, $n\sim 1000$ (worse case number for electronic bias of LNA and mixers)

GM RDK-3ST Sumitomo cryocooler maximal capacity on 4K stage



Cryogenic Machine ?

Cryocooler requirements to optimize receiver performances:

- Available power on 4K stage > 1.2 W (+ safety margin)
- Temperature on the *loaded* 4K stage < 4.2 K
- Stability: minimize temperature fluctuations of LNA !

Solutions?

- Use 2 cryocoolers ? (Cost, complexity, space required ...)
- Reduce power consumption of cryogenic LNA
- Not to operate LNA or SIS mixers at the optimal temperature

Requirements for cryostat:

- Low weight → aluminum
- Shape optimized to minimize receiver size and maximize ease of access to cryogenic components for repair or upgrade



RDK-3ST
Sumitomo
cryocooler